

Study of the Destruction of ^{18}F in Novae with an Inverse (d,p) Reaction at the HRIBF

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The degree to which the (p,γ) and (p,α) reactions destroy ^{18}F at temperatures $1\text{-}4\times 10^8$ K is important for understanding the synthesis of nuclei in nova explosions and for using the long-lived radionuclide ^{18}F , a target of γ -ray astronomy, as a diagnostic of nova mechanisms. The reactions are dominated by low-lying proton resonances near the $^{18}\text{F}+p$ threshold ($E_x=6.411$ MeV in ^{19}Ne). To gain further information about these resonances, we have used a radioactive ^{18}F beam from the Holifield Radioactive Ion Beam Facility (HRIBF) to selectively populate corresponding mirror states in ^{19}F via the inverse $^2\text{H}(^{18}\text{F},p)^{19}\text{F}$ neutron transfer reaction. The beam energy was 108.5 MeV (~ 6 MeV/u) and the target was a $160\ \mu\text{g}/\text{cm}^2$ CD_2 foil. Proton-recoil coincidence data were taken for both α -decaying and particle-stable final states using strip detector arrays and the Daresbury Recoil Separator (DRS) [1]. Neutron spectroscopic factors were measured for states in ^{19}F in the excitation energy range 0-9 MeV[2]. Widths for corresponding proton resonances in ^{19}Ne were calculated using a Woods-Saxon potential. The results imply a significant lowering of $^{18}\text{F}(p,\gamma)^{19}\text{Ne}$ and $^{18}\text{F}(p,\alpha)^{15}\text{O}$ reaction rates, thereby increasing the prospect of observing the 511-keV annihilation radiation associated with the decay of ^{18}F in the ashes ejected from novae[3]. Other recent and future experiments on the $A=19$ system at the HRIBF include measurements of proton resonances in ^{19}Ne via the $^2\text{H}(^{18}\text{F},^{15}\text{O}\alpha)n$ reaction[4] and additional measurements of the $^{18}\text{F}(p,\alpha)^{15}\text{O}$ excitation function[5]. Using a radioactive ^{26}Al beam at the HRIBF, techniques similar to those used in the $^2\text{H}(^{18}\text{F},p)^{19}\text{F}$ experiment will be employed for the $^2\text{H}(^{26}\text{Al},p)^{27}\text{Al}$ reaction to study the $^{26}\text{Al}(p,\gamma)^{27}\text{Si}$ reaction rate in stellar environments[6].

[1] R. Fitzgerald *et al.*, Nucl. Phys. **A748**, 351 (2005).

[2] R. L. Kozub *et al.*, Phys. Rev. C **73**, 044307 (2006).

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[4] C. R. Brune *et al.*, <http://www.phy.ornl.gov/hribf/news/feb-06/brune.shtml> (HRIBF Experiment RIB-145).

[5] K. Y. Chae *et al.*, Phys. Rev. C, in press (2006) and references therein; D. W. Bardayan *et al.*, HRIBF Experiment RIB-099.

[6] S. D. Pain *et al.*, HRIBF Experiment RIB-153.