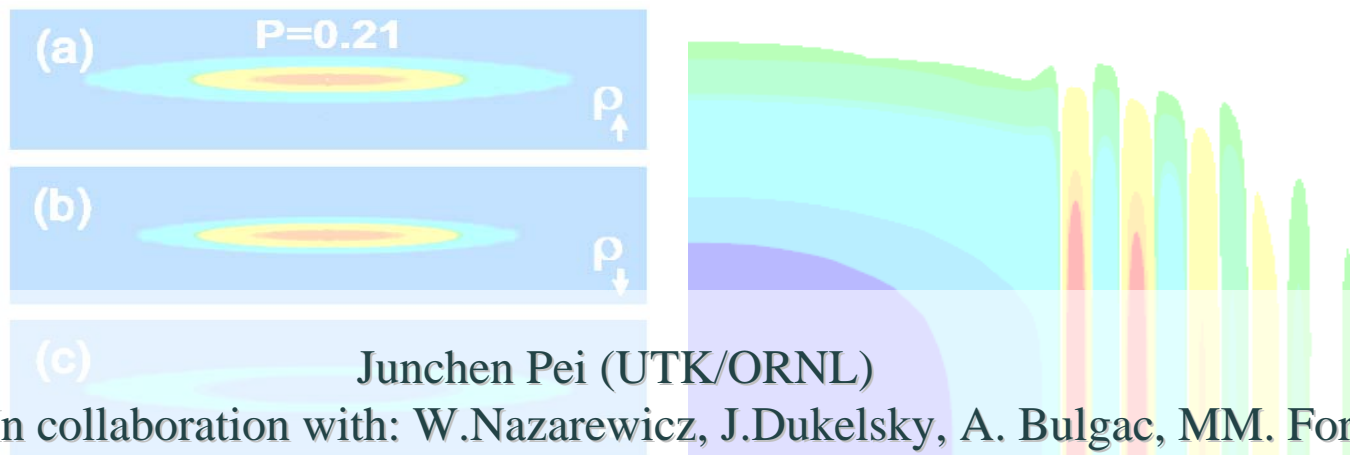


# Numerical Study of Exotic Pairing Phases in Imbalanced Fermi Condensates



# Pairing (superfluidity) in Fermi Systems

- ❑ Pairing (superfluidity is produced by forming Cooper pairs of two spin components) is a generic feature in strongly interacting Fermi systems
- ❑ Pairing phases are more interesting in imbalanced systems: with mismatched chemical potentials of the two spin states.

## Imbalanced (or polarized or asymmetric) systems:

- magnetized superconductors
- Cold dense quark matter in compact stars :  $\delta\mu = \frac{1}{2}(\mu_d - \mu_u) = \frac{1}{2}\mu_e$   
(R.Casalbuoni, et al, RMP 76,263, 2004)
- ultracold atomic gas: two hyperfine states of  ${}^6\text{Li}$
- Nuclear physics (odd-nuclei, neutron-skin, n-p pairing, neutron stars, asymmetric nuclear matter)

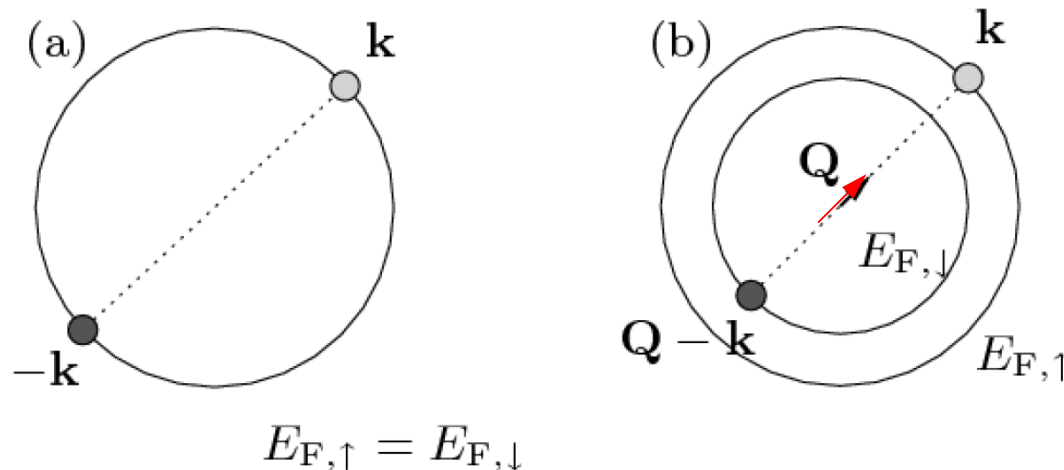
## Exotic pairing phases (temperature and polarization):

- Gapless excitations. (also happen in nuclear high-spin states: G.Bertsch et al, PRA 79, 043602 (2009) )
- Phase separation: a paired superfluid core surrounded by an unpaired shell  
G.B.Partridge, et al., Science 311, 503
- FFLO pairing phase

# Expected exotic FFLO pairing

- In imbalanced Fermi systems, pairing with non-zero momentum can happen: Flude-Ferrell-Larkin-Ovchinnikov (**FFLO**)  
Oscillation pairing gap is expected; Modulated densities (crystallized).
- It exists in many theoretical calculations, but difficult to find.
- Some signatures in heavy fermions systems. [Radovan, et al. Nature 425, 51, 2003.](#)

$$\Delta(\vec{x}) = \Delta_0 \sum_n C_n e^{i\vec{q}_n \cdot \vec{x}}$$



# Some experiments

- Advantages of using cold atoms:  
interaction is controllable; clear physics;  
High  $T_c$ ; implications for other Fermi systems
- Unitary limit:** two body s-wave scattering length diverges:  $a_s \rightarrow \pm \infty$

System is strongly correlated and its properties do not depend on the value of scattering length  $a_s$

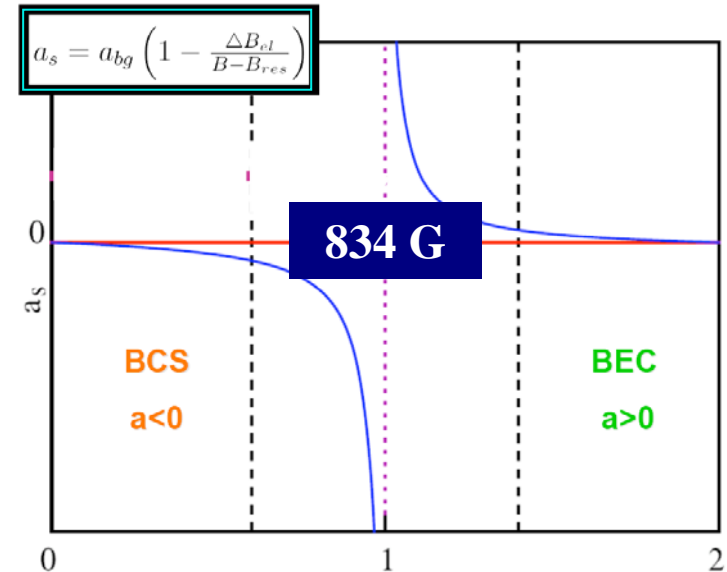
- It is realized through Feshbach Resonance: magnetic field at 834 G
- Trapped by optical and magnetic potential

$$U(r, z) = \frac{1}{2} m \omega_B^2 z^2 + U_0 \left[ 1 - \frac{w_0^2}{w^2(z)} e^{-2r^2/w^2(z)} \right], \quad w(z) = w_0 \left[ 1 - (z/z_0)^2 \right]$$

approximate HO potential:  $U(r, z) = U_0 \left( 1 - \exp\left(-\frac{1}{2} (w_r^2 r^2 + w_z^2 z^2) / U_0\right) \right)$

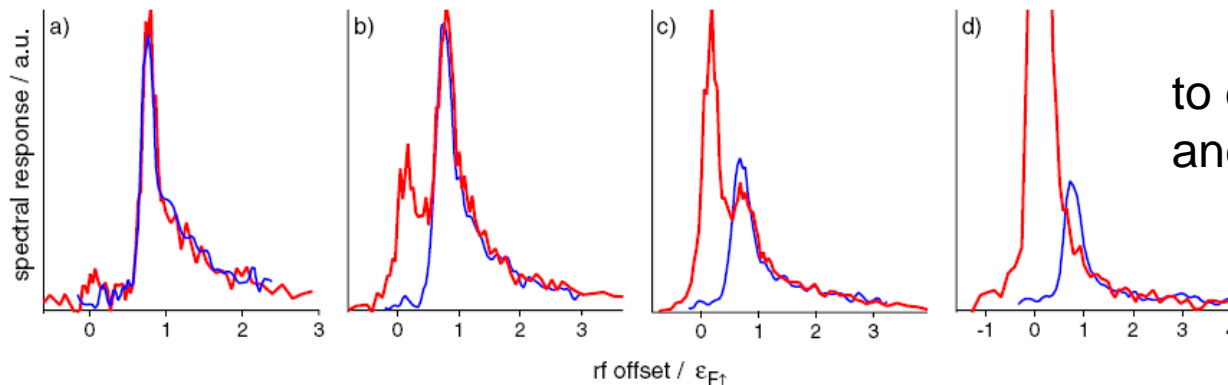
**Aspect ratio:**  $\eta = w_r / w_z$

highly elongate trap is of great interests! (good for looking FFLO pairing)



# Some experiments

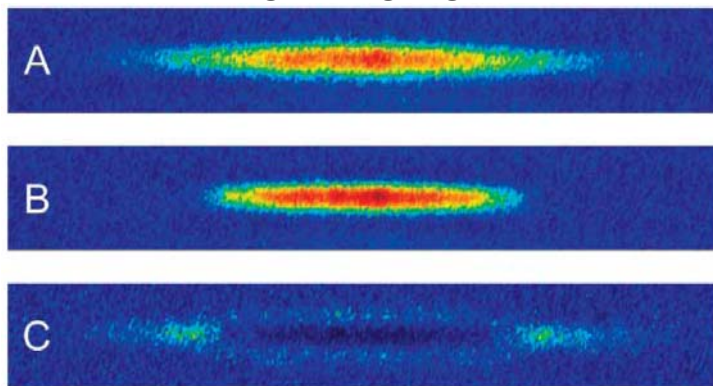
- Radio frequency spectrum



to determine the gap  
and Hartree energy

Schirotzek, et al. PRL101,140403.

- *In situ* absorbing imaging



Spin-up density

Spin-down density

polarized density

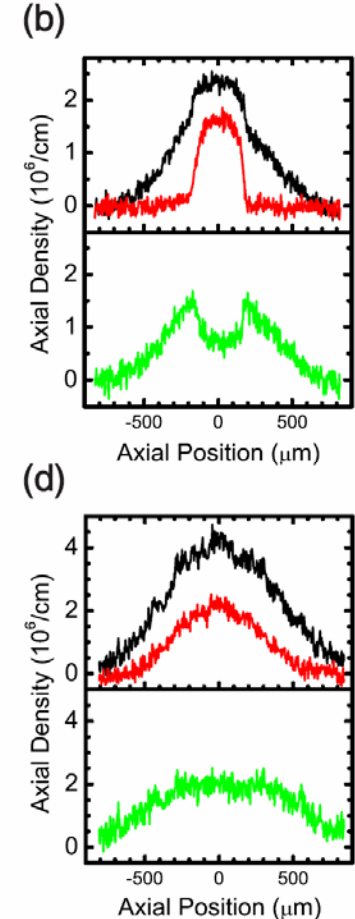
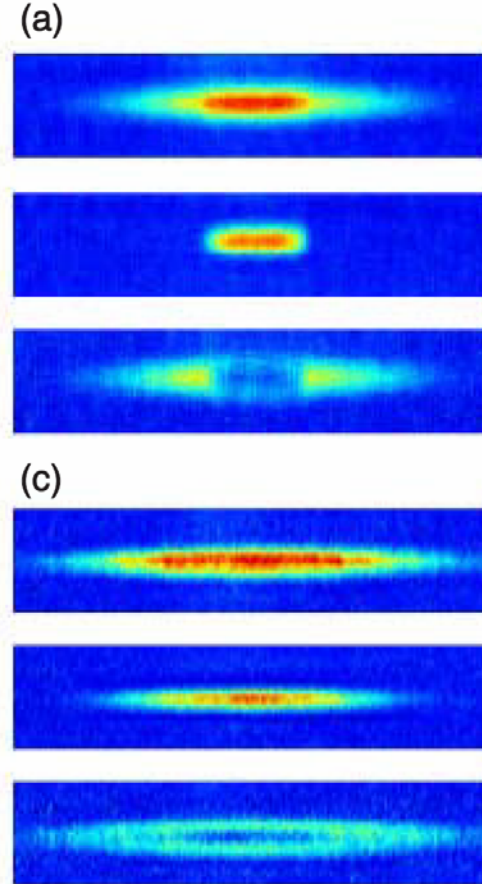
Science 311(2006)503

# Experiments(Rice)

- Phase Separation
- Superfluid Core is deformed from the trap shape and such deformation effects disappear at high temperatures
- Trap aspect ratio  $\sim 50$ : highly elongate
- Particle numbers  $\sim 10^5$

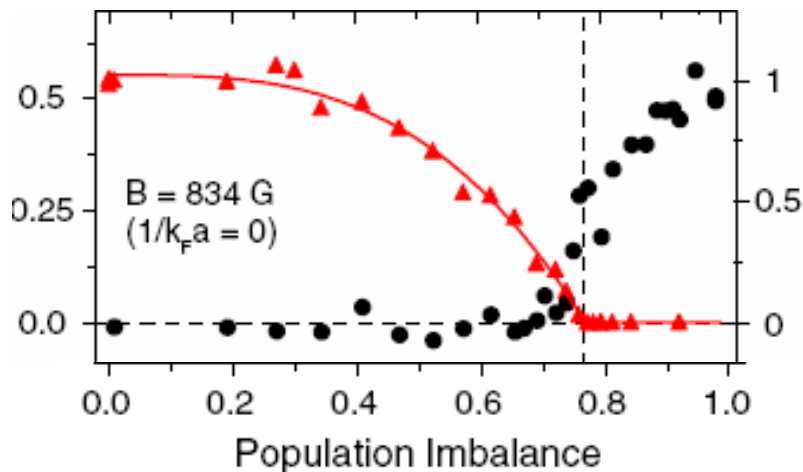
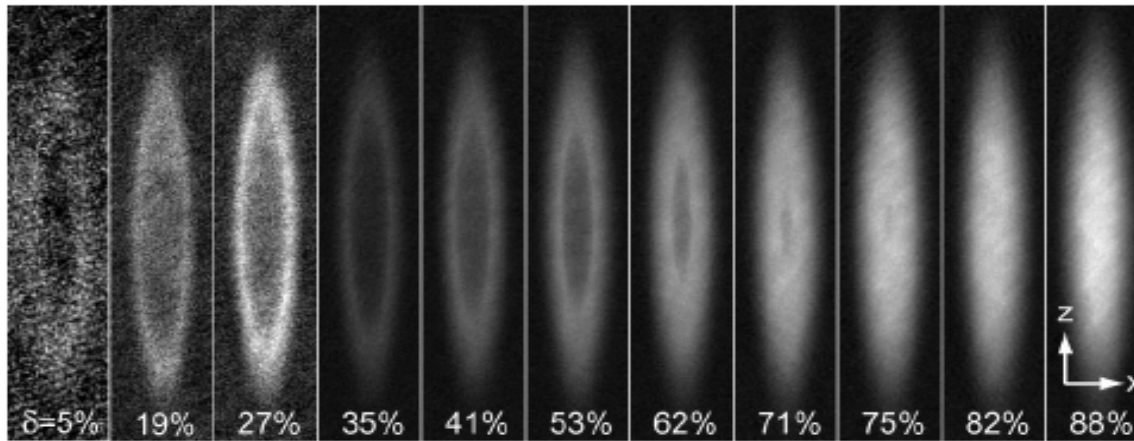
G.B.Partridge, et al, PRL97,190407,2006

G.B.Partridge, et al, Science,311,503,2006



# Experiments(MIT)

- Phase separation
- However, no superfluid core deformation Y.Shin, et al, PRL 97,03401,2006

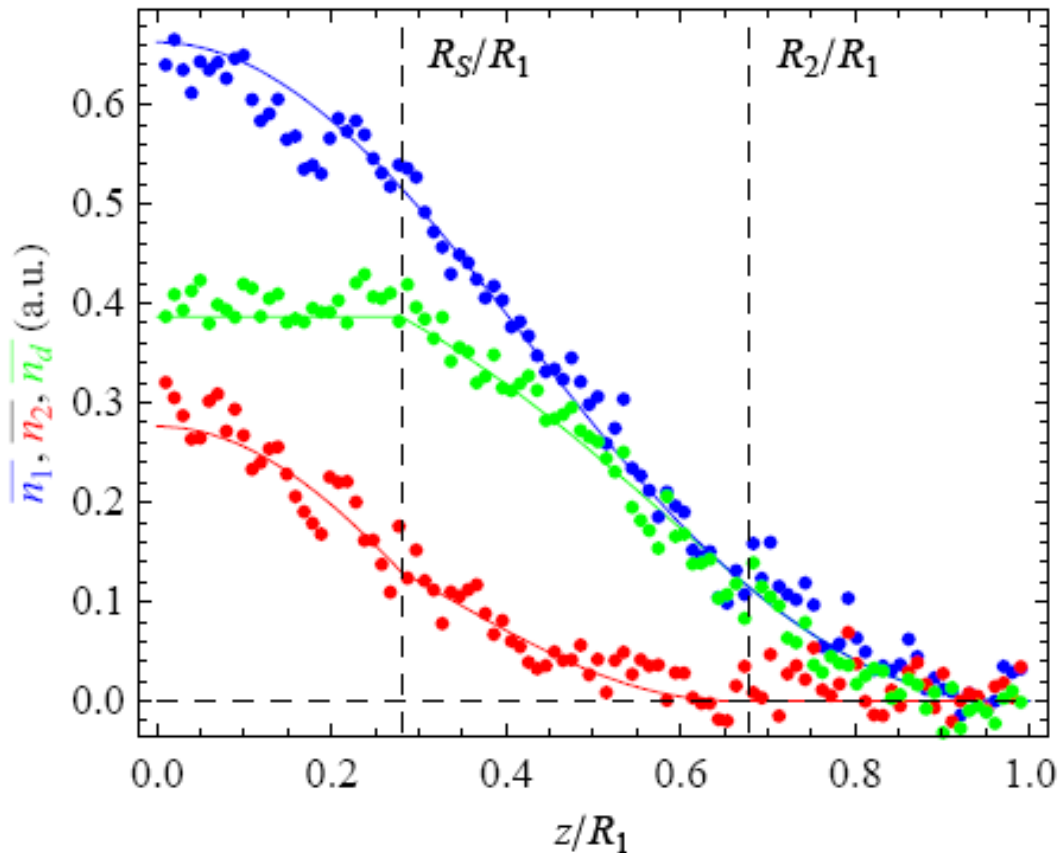


- Clogston-Chandrasekhar limit of superfluidity
- Trap aspect ratio=5, particles= $10^6$

# Experiments-others

- French group:  $10^5$  particles, aspect ratio=23 (agree with MIT)

No core deformation



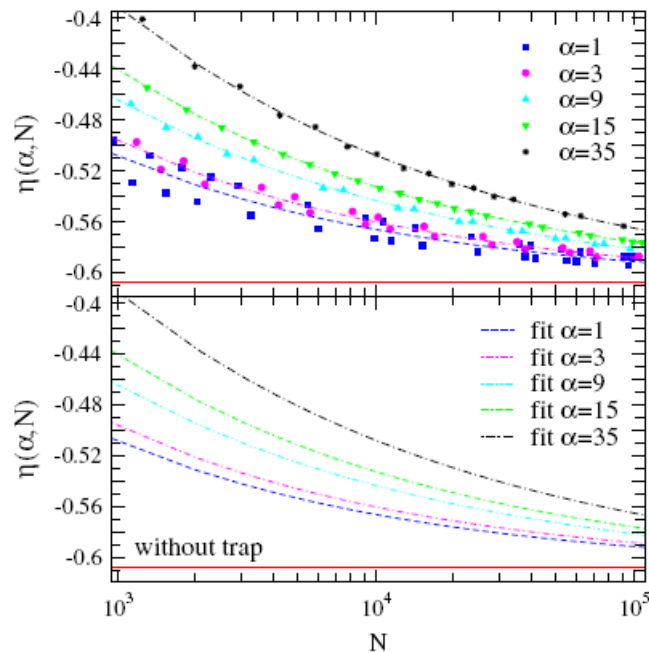
C. Salomon, et al, PRL103, 18 (2009) 170402

**Question???**

1. different experimental conditions
2. or theory is not precise

# Finite-size effect

- Finite-size effect of trap deformations and particle numbers  
small deformation trap doesn't violate LDA solutions



M.Ku, PRL 102, 255301, 2009

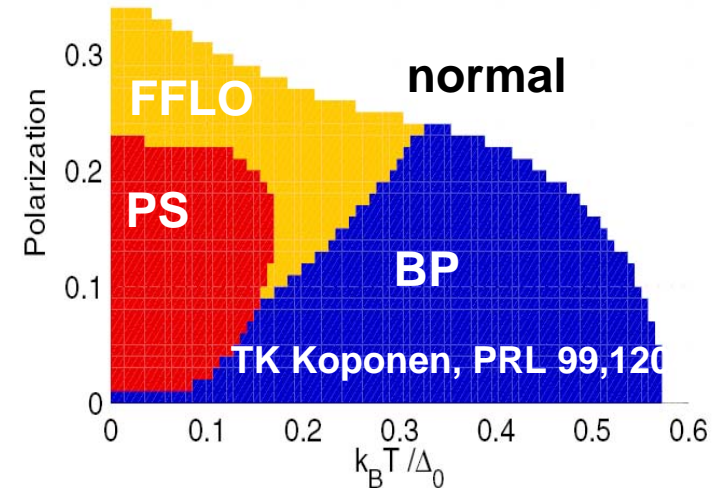
- Surface tension is important at large deformations  
T.N. De Silva, et al, PRL 97, 070402(2006)

# Theoretics

- Quantum Monte Carlo: QMC is very precise but limited to small systems
- Bogoliubov de-Genes equation: Mean Field approximation

$$H(\mathbf{r}) = \sum_{\sigma} \Psi_{\sigma}^{\dagger}(\mathbf{r}) [H_0(\mathbf{r}) - \mu_{\sigma}] \Psi_{\sigma}(\mathbf{r}) - g \Psi_{\uparrow}^{\dagger}(\mathbf{r}) \Psi_{\downarrow}^{\dagger}(\mathbf{r}) \Psi_{\downarrow}(\mathbf{r}) \Psi_{\uparrow}(\mathbf{r})$$

$$\begin{pmatrix} H_0(\mathbf{r}) - \mu & \Delta(\mathbf{r}) \\ \Delta^*(\mathbf{r}) & -H_0(\mathbf{r}) + \mu \end{pmatrix} \begin{pmatrix} u_i(\mathbf{r}) \\ v_i(\mathbf{r}) \end{pmatrix} = E_i \begin{pmatrix} u_i(\mathbf{r}) \\ v_i(\mathbf{r}) \end{pmatrix}$$



Plenty of calculations, no Hartree potential, and not quantitatively accurate  
 A contest of computation: Tokyo U: 30000 particles; Rice U:  $10^5$  particles

- DFT: at the unitary limit, the physical properties only depends on the density. It is good for DFT descriptions. Superfluid Local Density Approximation (SLDA) is very precise.

- ASLDA Equations:

$$\begin{bmatrix} h_a(\mathbf{r}) - \lambda_a & \Delta(\mathbf{r}) \\ \Delta^*(\mathbf{r}) & -h_b(\mathbf{r}) + \lambda_b \end{bmatrix} \begin{bmatrix} u_i(\mathbf{r}) \\ v_i(\mathbf{r}) \end{bmatrix} = E_i \begin{bmatrix} u_i(\mathbf{r}) \\ v_i(\mathbf{r}) \end{bmatrix}$$

$$\rho_a(\mathbf{r}) = \sum_i f_i |u_i(\mathbf{r})|^2, \quad \rho_b(\mathbf{r}) = \sum_i (1 - f_i) |v_i(\mathbf{r})|^2$$

$$f_i = 1 / (1 + \exp(E_i / kT))$$

$$\kappa(\mathbf{r}) = \sum_i f_i u_i(\mathbf{r}) v_i^*(\mathbf{r}), \quad \Delta(\mathbf{r}) = g_{eff}(\mathbf{r}) \kappa(\mathbf{r})$$

ASLDA energy density functional:

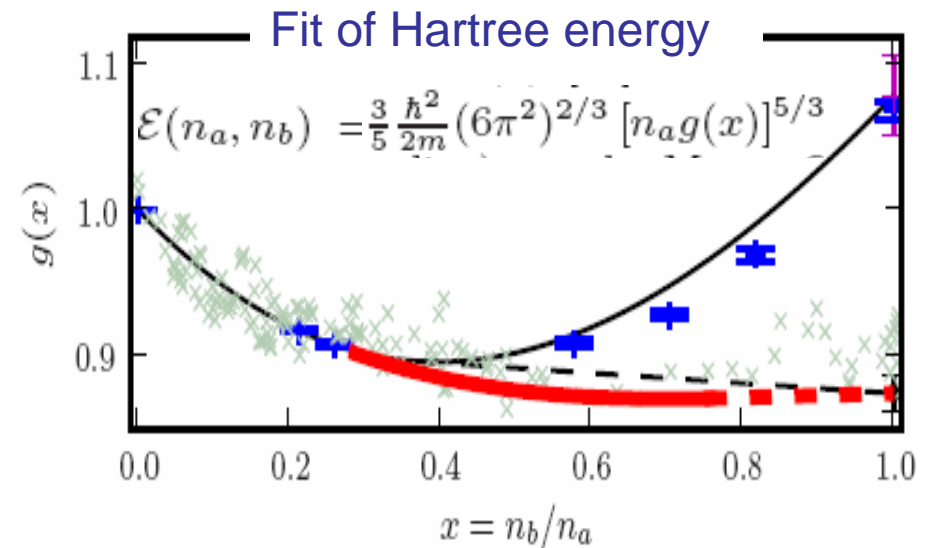
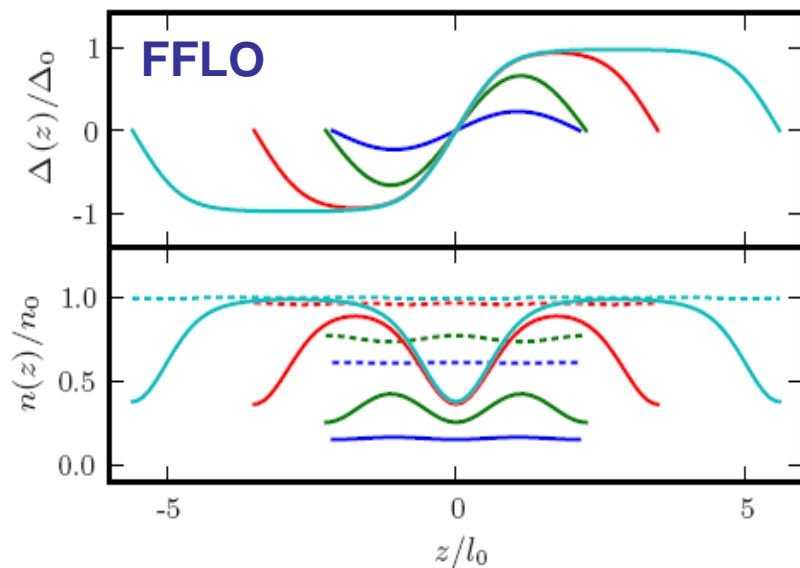
$$\mathcal{E} = \alpha_a(x) \frac{\tau_a}{2} + \alpha_b(x) \frac{\tau_b}{2} + \frac{(3\pi^2(\rho_a + \rho_b))^{5/3}}{10\pi^2} \beta(x) - \Delta \kappa$$

$$g_{eff}(\mathbf{r}) = \frac{\gamma(x)}{(\rho_a + \rho_b)^{1/3} + \Lambda(\mathbf{r})\gamma(x)}$$

Pairing regularization

- Becomes SLDA when effective mass of spin-up and spin-down is the same.

- Parameters are fitted according to experiments and QMC calculations (effective mass; energy density) as functions of polarizations close to SLDA at small polarizations
- ASLDA predicted FFLO phase: periodic boundary condition  
A.Bulgac, M.Forbes, PRL101:215301,2008



# Coordinate-space HFB calculations

- It is the advantage of coordinate-space HFB to treat elongated potential.  
**HFB-AX** code: using B-spline techniques; Axially symmetry; Very precise for deformed and weakly bound nuclei.

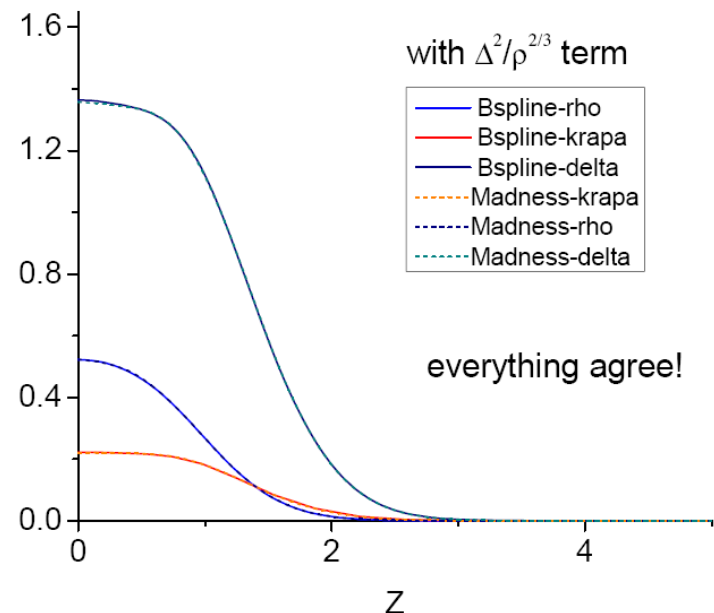
1000 atoms employed.

dimensionless 2D-box:  $R=25 \times 0.3$ ,  $Z=(125 \sim 300) \times 0.3$  (very dense spectrum)

the elongation is set to be  $\eta=10, 20, 30, 40$ . ASLDA is very time consuming

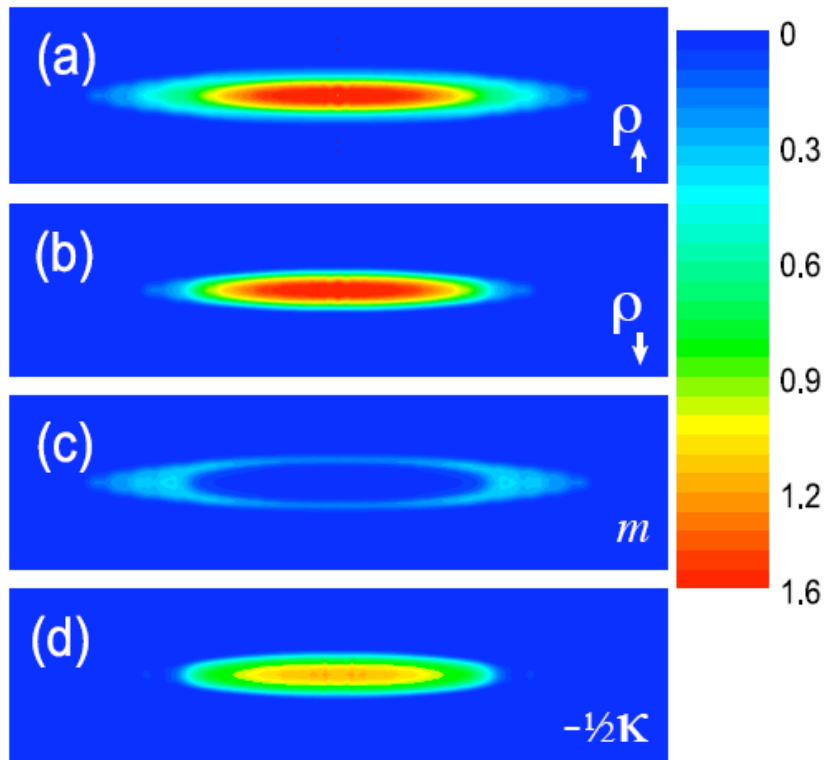
*JP et al, PRC 78, 064306, 2007*

- 3D HFB-Madness is under development (multi-resolution adaptive wavelet basis):  
SLDA-Madness has been benchmarked.  
ASLDA-Madness is being debug.



# Phase separation

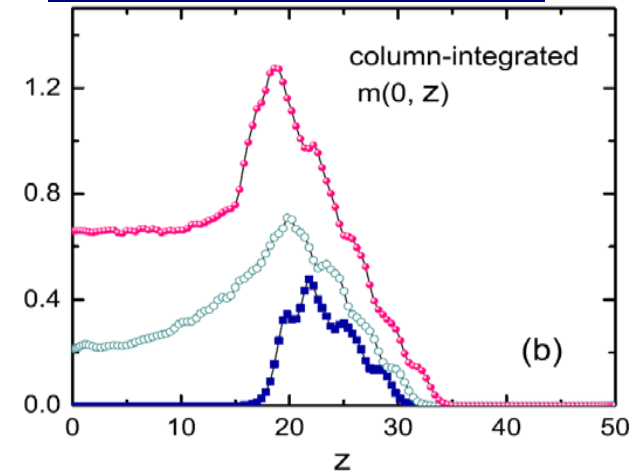
- SLDA calculations ( $\eta = 10$ ):



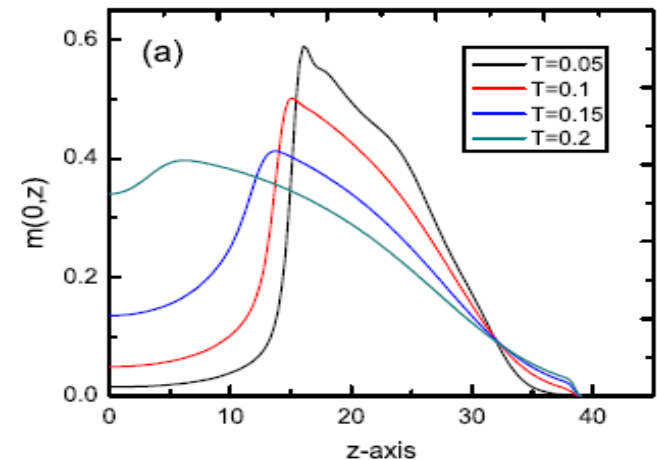
$$P = (N_a - N_b) / (N_a + N_b)$$

JP, W. Nazarewicz, M. Stoitsov, EPJA42,595 2009

## Polarization effects

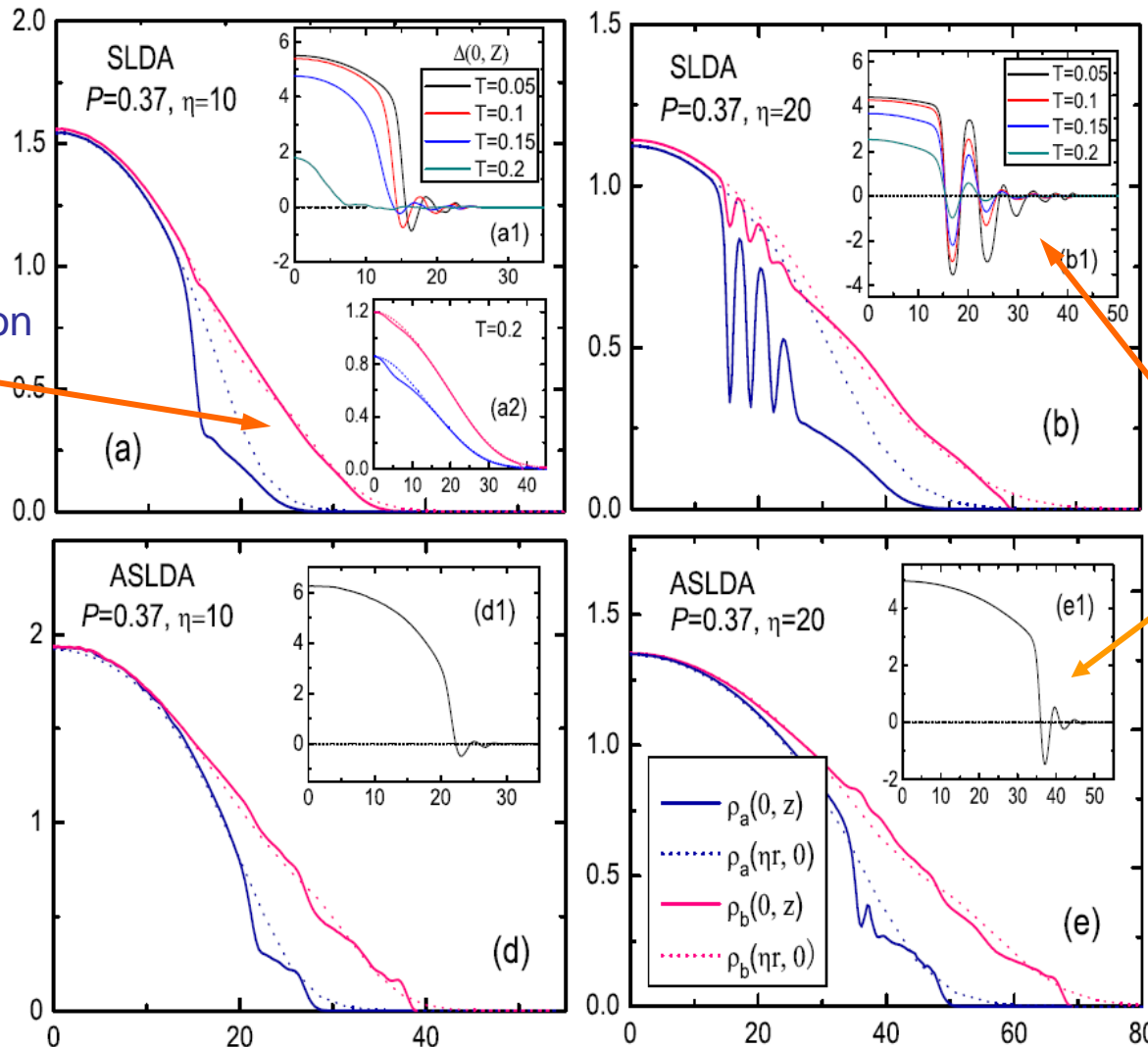


## Temperature effects



# Traps with small elongation

- Core deformation is predicted in a small deformed trap by SLDA

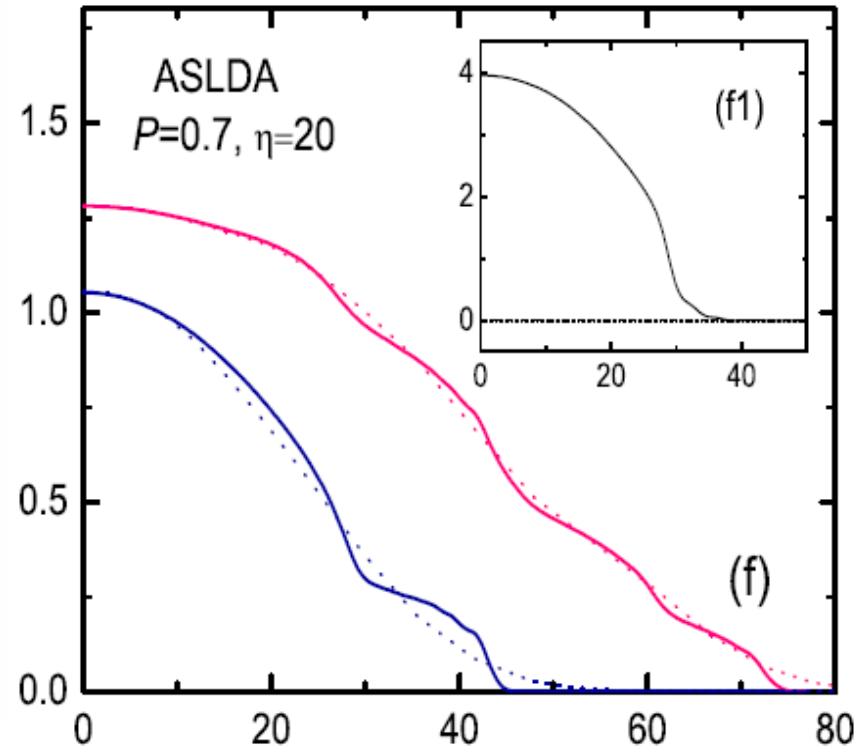
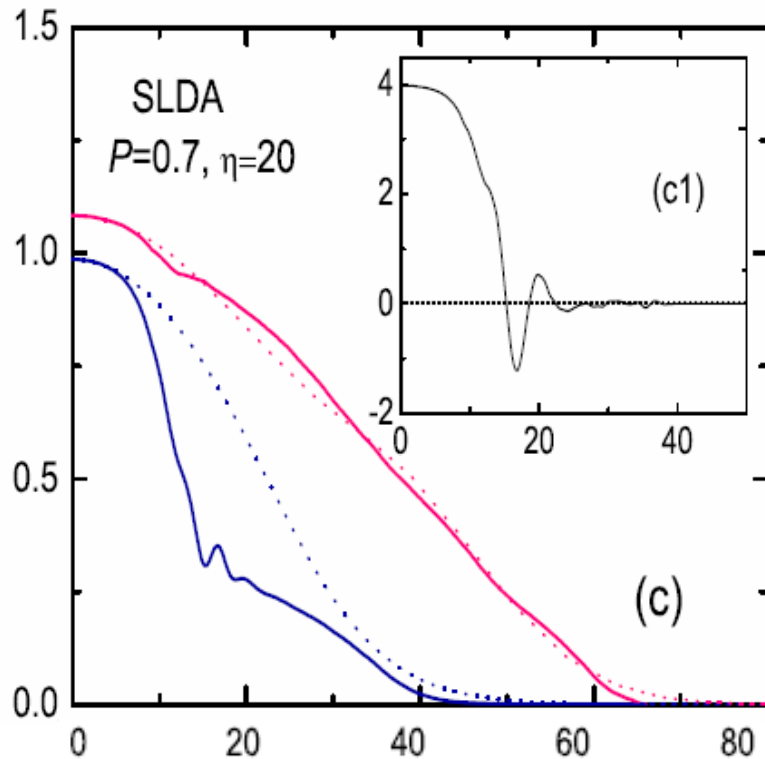


Core deformation

FFLO

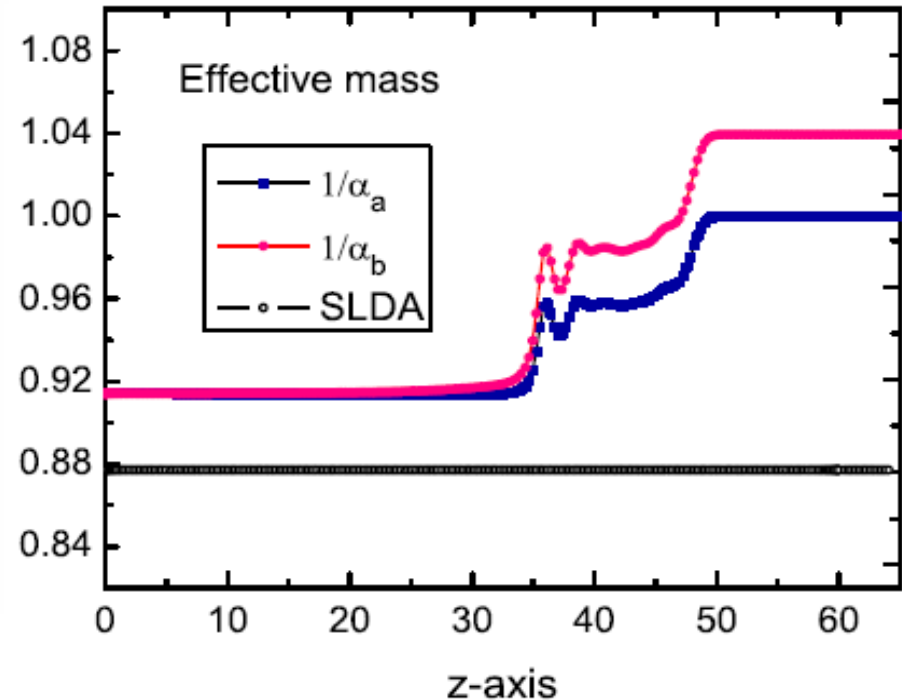
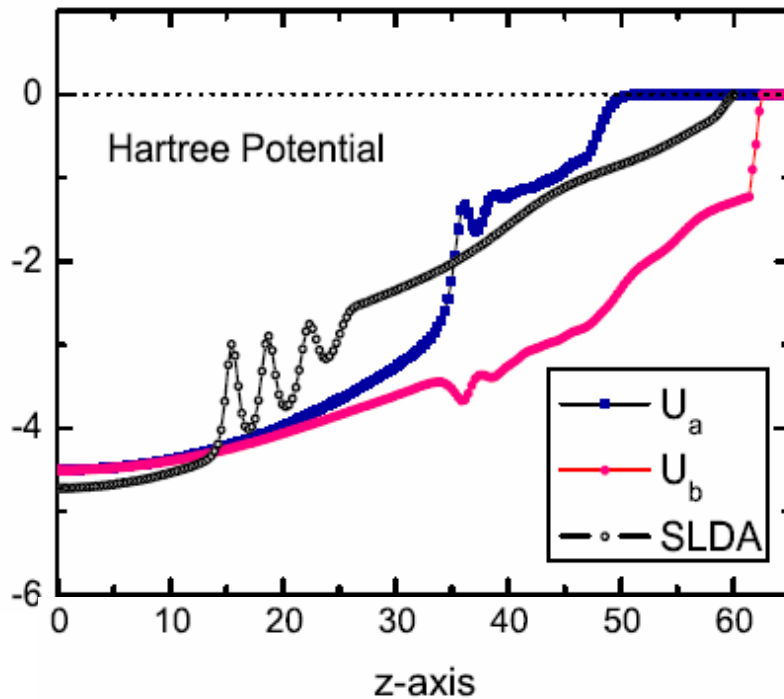
# Large polarization

- FFLO and phase separation fades away
- ASLDA predicts less significant FFLO than SLDA



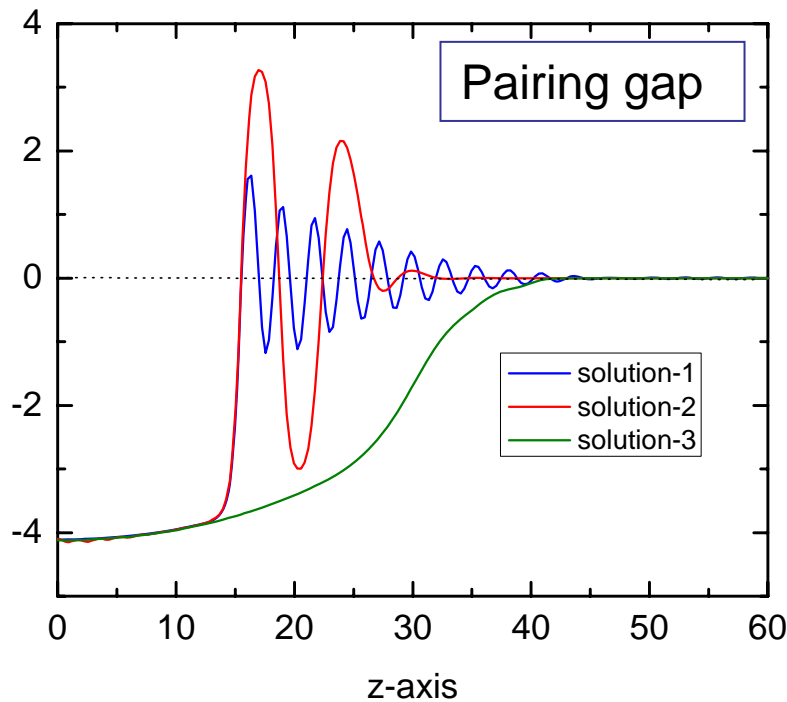
# SLDA and ASLDA difference

- Very different Hartree potentials in ASLDA
- Different effective mass



# Multi-solutions in SLDA

- Different initial conditions lead to different solutions(  $\eta = 20$ ).  
**solution3**- start from flat density distributions  
**solution2**- start from less deformed solution  $\eta = 10$   
**solution1**- start from 2 times frequency of solution-2 (artificial)

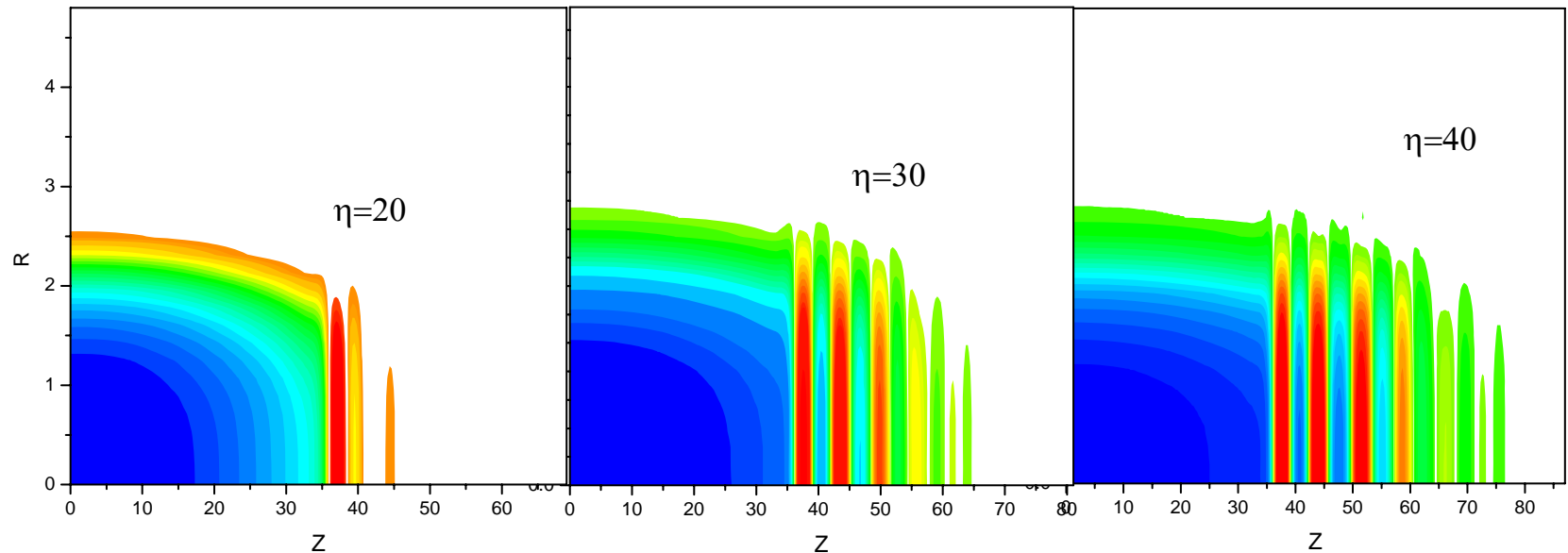


solutions	energy
<b>1</b>	3250
<b>2</b>	3305
<b>3(G.S)</b>	3150

Oscillation amplitudes decrease as frequency increase

# Increase the trap aspect ratio

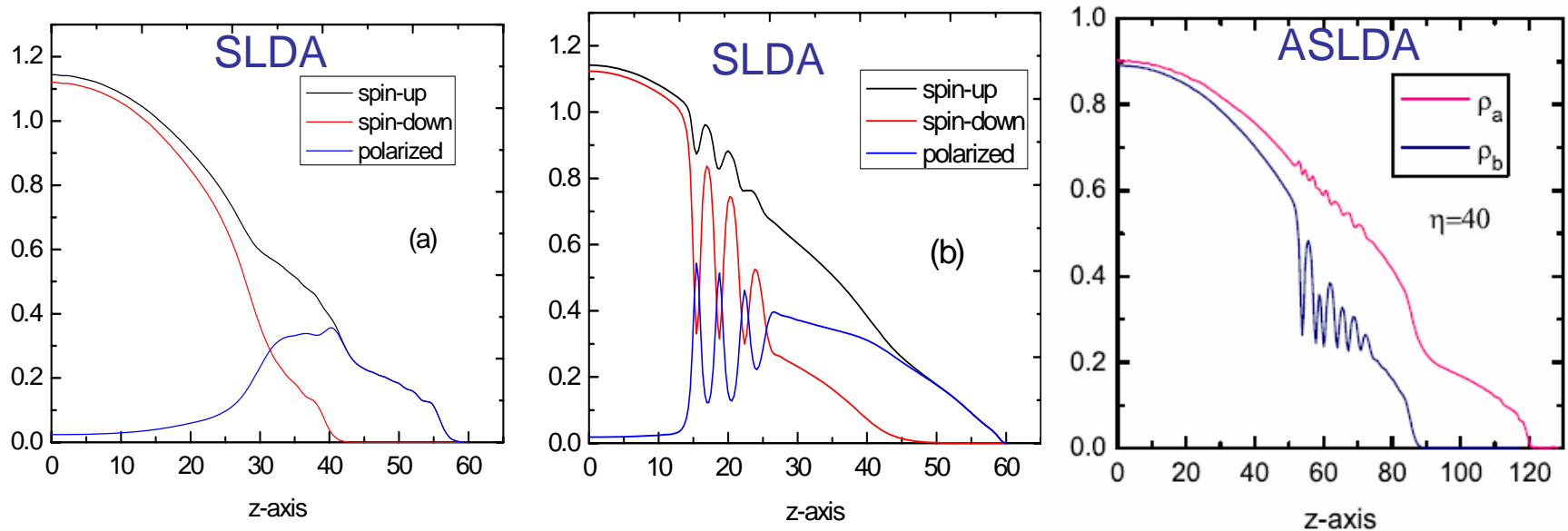
- Pairing oscillations become remarkable as trap aspect ratio increase
- The oscillations are perpendicular to the long axis
- Oscillation periods are the same



ASLDA calculations of polarization=0.37

# FFLO detection

- Different densities of the LDA and FFLO solutions should be distinguished by *in situ* imaging.



- How to access the FFLO in experiments (excitation state): we suggest increase the trap elongation after the condensation obtained

# Conclusions

- Coordinate-space HFB is good for Fermi condensates study.
- Phase separation is demonstrated in both SLDA and ASLDA
- Phase separation disappears as temperature becomes higher or polarization becomes larger.
- Superfluid Core deformation is not shown in ASLDA. Still ambiguous
- The FFLO is predicted in highly elongated trap, both by SLDA and ASLDA. However, It has a higher energy than LDA solution. Difficult for experiments to access. Calculation still continues.
- A future fully precise DFT study of  $10^5$  particles is very valuable