

Natural Units in the Nuclear Energy Density Functional Theory

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Skyrme EDF

- Most used EDF in nuclear physics

$$E_t^{\text{even}} = C_t^{\rho} \rho_t^2 + C_t^{\tau} \rho_t \tau_t + C_t^{\Delta\rho} \rho_t \Delta\rho_t^2 + C_t^{\nabla J} \rho_t \nabla J_t + C_t^J J_t^2, \quad t=0,1$$

$$C_t^{\rho} = C_{t0}^{\rho} + C_{tD}^{\rho} \rho_0^{\sigma}$$

- Uses local density approximation:

$$V(\mathbf{r}_1, \mathbf{r}_2) \propto \delta(\mathbf{r}_1 - \mathbf{r}_2) \Rightarrow \rho(\mathbf{r}, \mathbf{r}') \rightarrow \rho(\mathbf{r}, \mathbf{r}) = \rho(\mathbf{r})$$

- Density dependence included in ρ_t^2 term
- Derivatives of densities up to second order

Effective Lagrangian

- Each term in effective low-energy Lagrangian is written as

$$g \left[\frac{\psi \psi^\dagger}{\Lambda f_\pi^2} \right]^l \left[\frac{\nabla}{\Lambda} \right]^n \Lambda^2 f_\pi^2$$

- f_π^2 is pion-decay constant, g is a dimensionless constant close to unity and Λ characterizes energy scale beyond effective Lagrangian
- Standard Skyrme coupling constants are scaled by a factor of

$$S = f_\pi^{2(l-1)} \Lambda^{n+l-2}$$

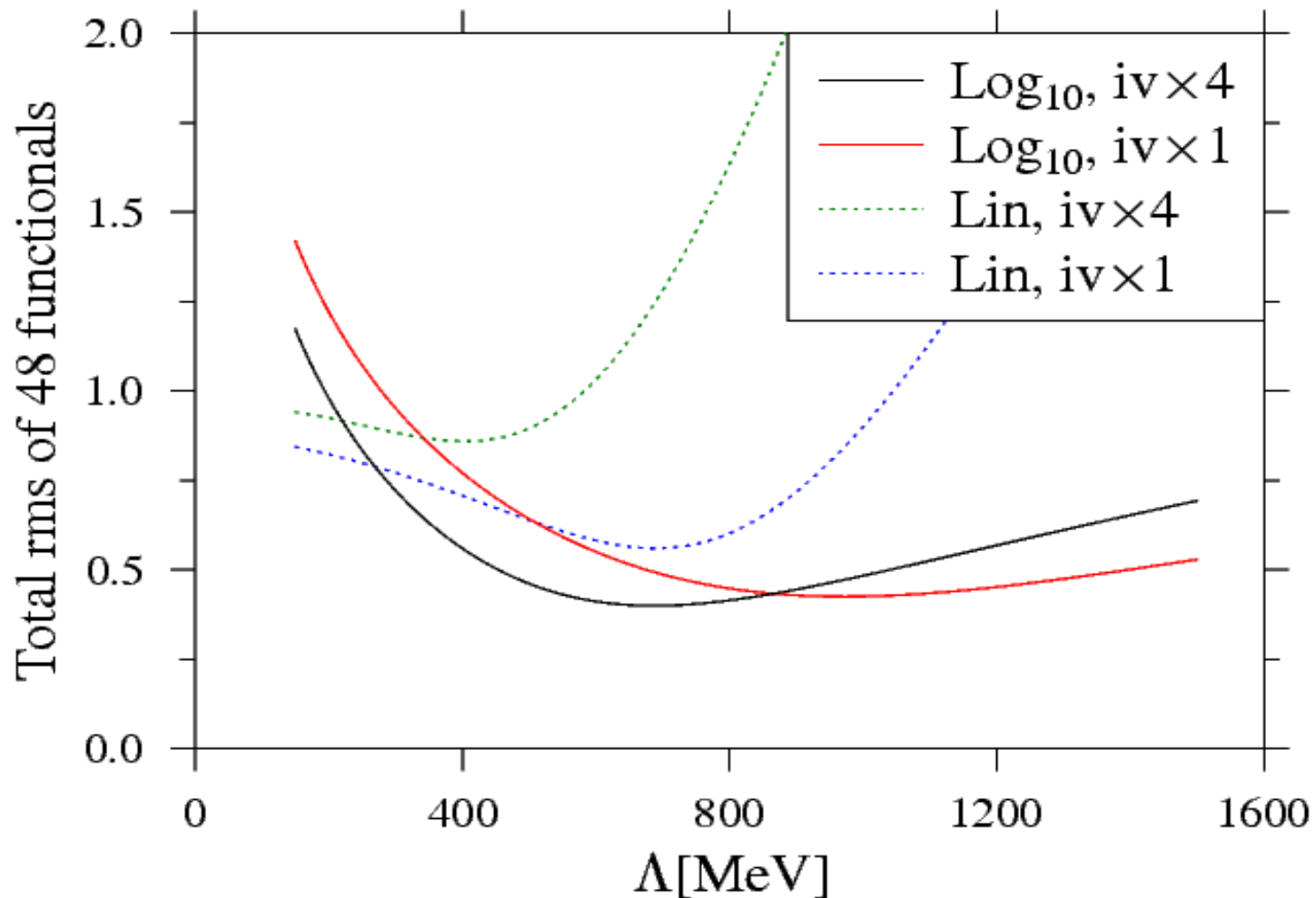
- Isovector coupling constants in natural units are scaled by a factor of 4
- Continuation of Furnstahl & Hackworth, PRC56, 2875
- Natural units scaling applied to 48 Skyrme functionals

List of functionals

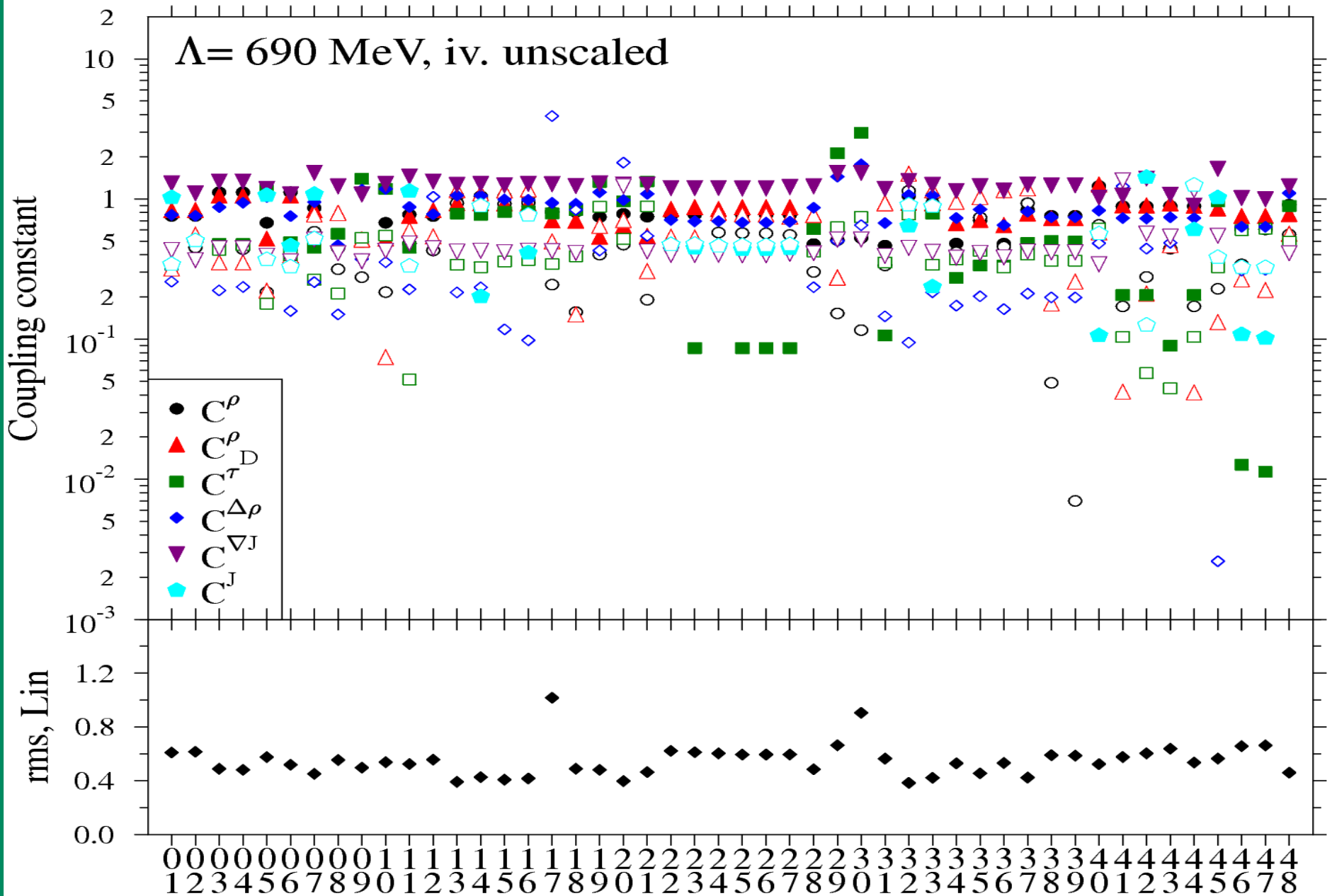
1—2	SkT3, SkT6	a d i j	a) masses of double-magic nuclei
3	SkM	a c e f g i	(includes ^{90}Zr , ^{116}Sn , ^{124}Sn and ^{140}Ce)
4	SkM*	g j	used in the fit
5—6	SGI, SGII	d e j	b) masses of non-double-magic nuclei
7	HFB9	a b f h i	used in the fit
8—9	SI, SII	a c d e f i	c) charge radii used in the fit
10	SkA	a c d e g i j	d) single particle energies used in the fit
11	HFB16	a b c f h i	e) symmetric infinite nuclear matter
12	SkT	a b d e g h i	constrains considered in the fit
13—16	Sly4—7	a c d e f i	f) asymmetric infinite nuclear matter
17—18	SkI1—2	a b c d f g i	constrains considered in the fit
19—20	SkI3—4	a b c d f g i k	g) surface properties (neutron skin,
21	SkI5	a b c d f g i	fission barriers etc.) considered in the
22—27	Msk1—6	a b f h i	fit
28—29	SIII, SIV	a c i	h) pairing was present in the fit
30—31	SV, SVI	a c i j	i) some parameters were fixed in the fit
32—33	Sly230a, b	a c d e f i	j) parameters extrapolated or fine-tuned
34—39	E, E_{σ} , Z, Z_{σ} , R_{σ} , G_{σ}	a c d g i	from existing force or functional
40	SkP	a b c e f h i	k) functional is an extended functional
41—42	SkO, SkO'	a b c d f g i k	
43	SV-min	a b c d g h k	
44	SkO _T	i j k	
45	SkMP	a j	
46—47	SkX, SkX _c	a b c d e f	
48	RATP	a d e f i	

Choice of optimal Λ

- One way to obtain optimal Λ is to minimize the deviation of coupling constants from unity \Rightarrow rms fit



Coupling constants in natural units

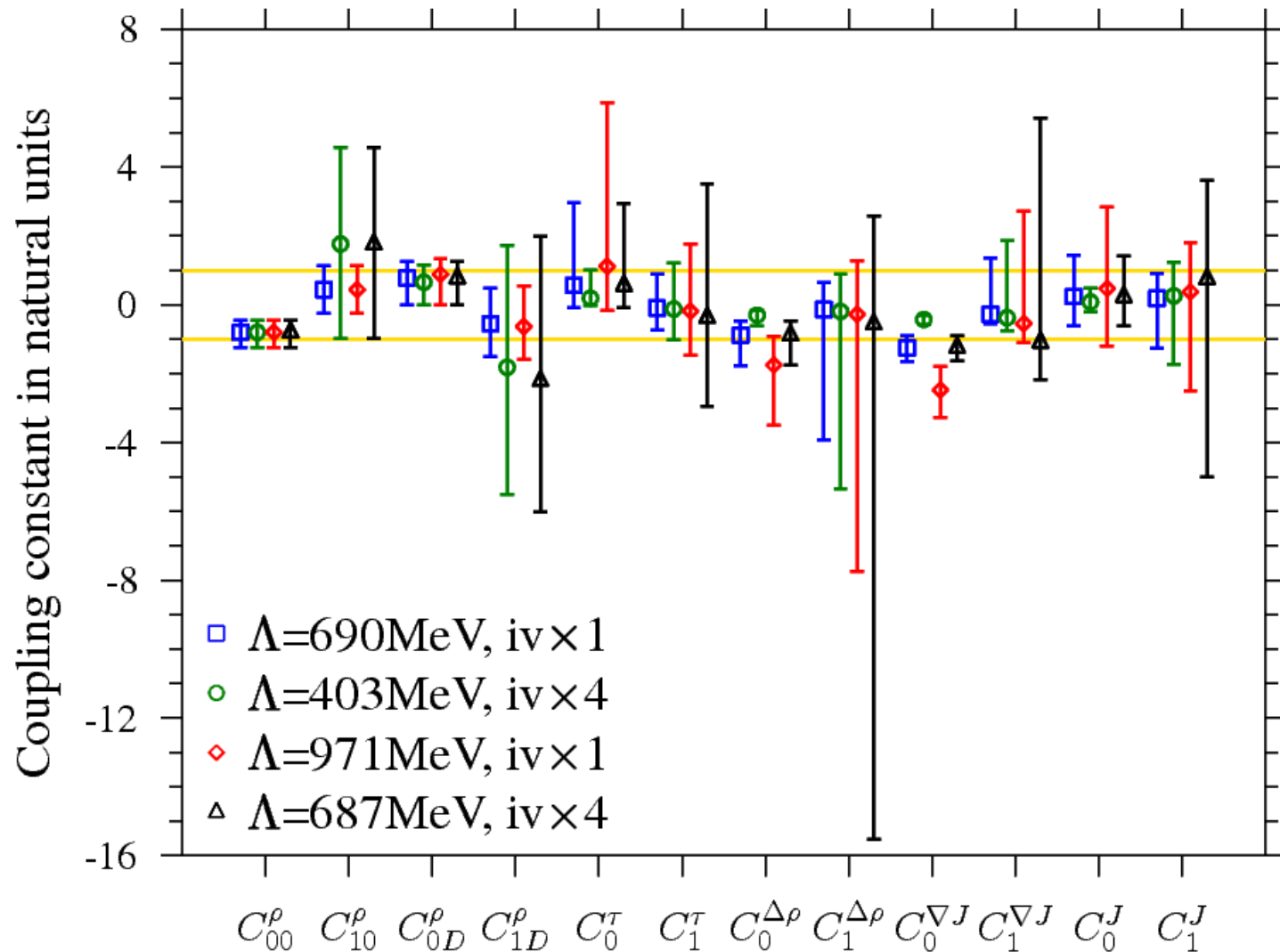


Deviation from unity

- Deviation from unity linked in many cases to uncomplete optimization procedure
- Anomalously small $C^{\Delta\rho}_1$ in RATP explained by a focus on volume part of the functional for astrophysical purposes
- Small $C^{\Delta\rho}_1$ in SkMP comes from mixing SkM* and SkP functionals with only minor adjustment to volume part
- Large $C^{\Delta\rho}_1$ in SkI1 compared to other SkI functionals due to the lack of isotopic shift data in optimization
- Large C^τ_0 in SV due to artificially imposed zero density dependence
- Small C^τ_0 in SkX results from strong emphasis on single particle data which favors effective mass close to one

Range of coupling constants in natural units

- Naturalness can be used as a guide in future fitting attempts
- Range and average of cc's in 48 functionals:



Conclusions

- Scaling to natural units brings coupling constants close to unity
- Parameter Λ around expected value for optimal rms
- Large deviations from unity can be linked to uncomplete optimization procedure in many cases
- Natural units can be used as a guiding principle in future fitting attempts
- Natural units also included in massexplorer.org

People involved

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